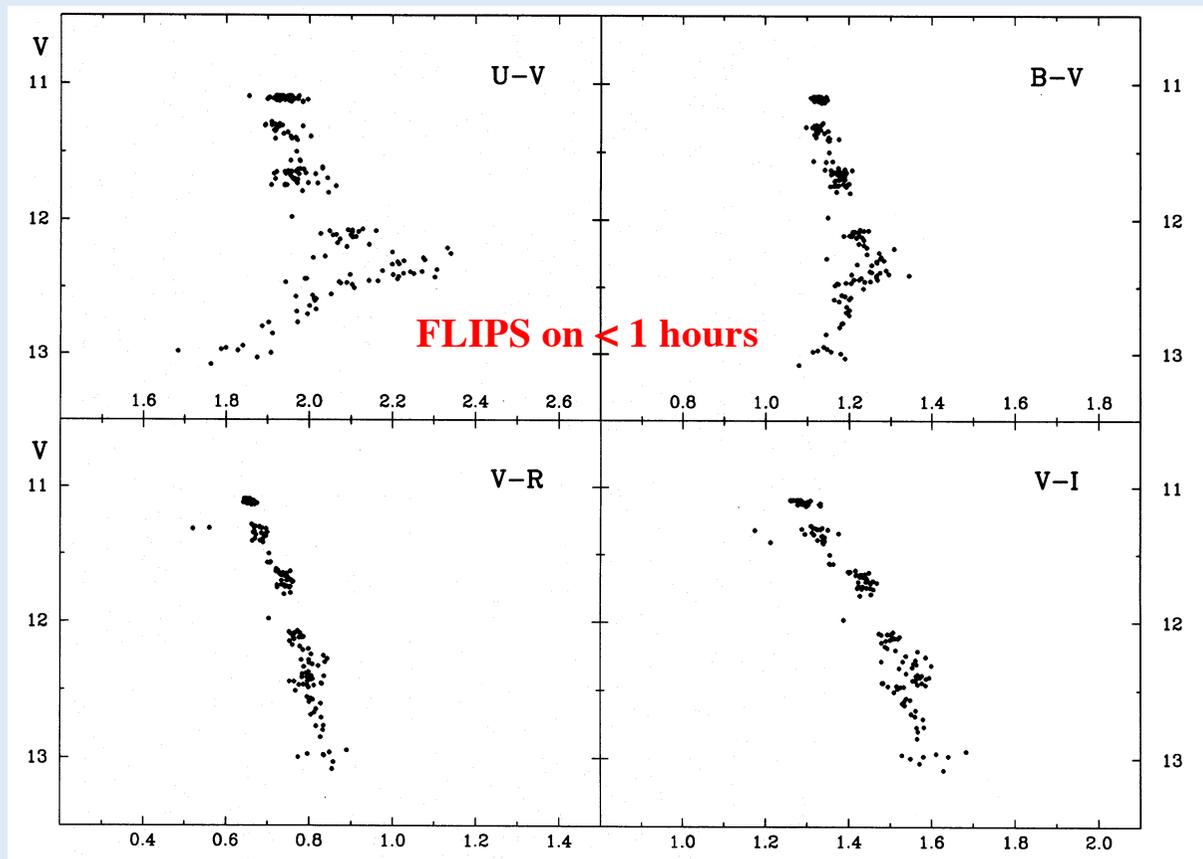
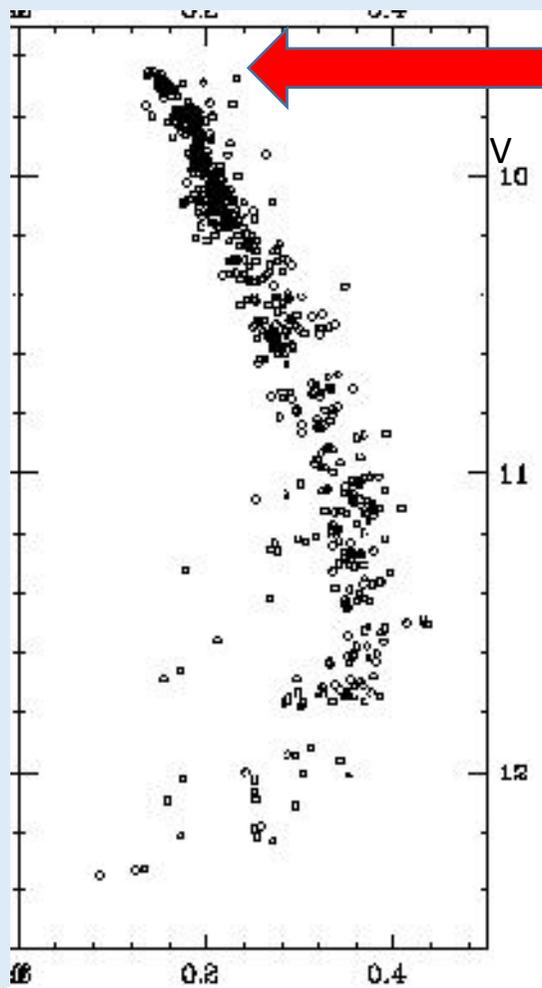


Covered by large dust grains in the foreground disk?
Gahm + (1988, 1993): RY Lup, type TTS



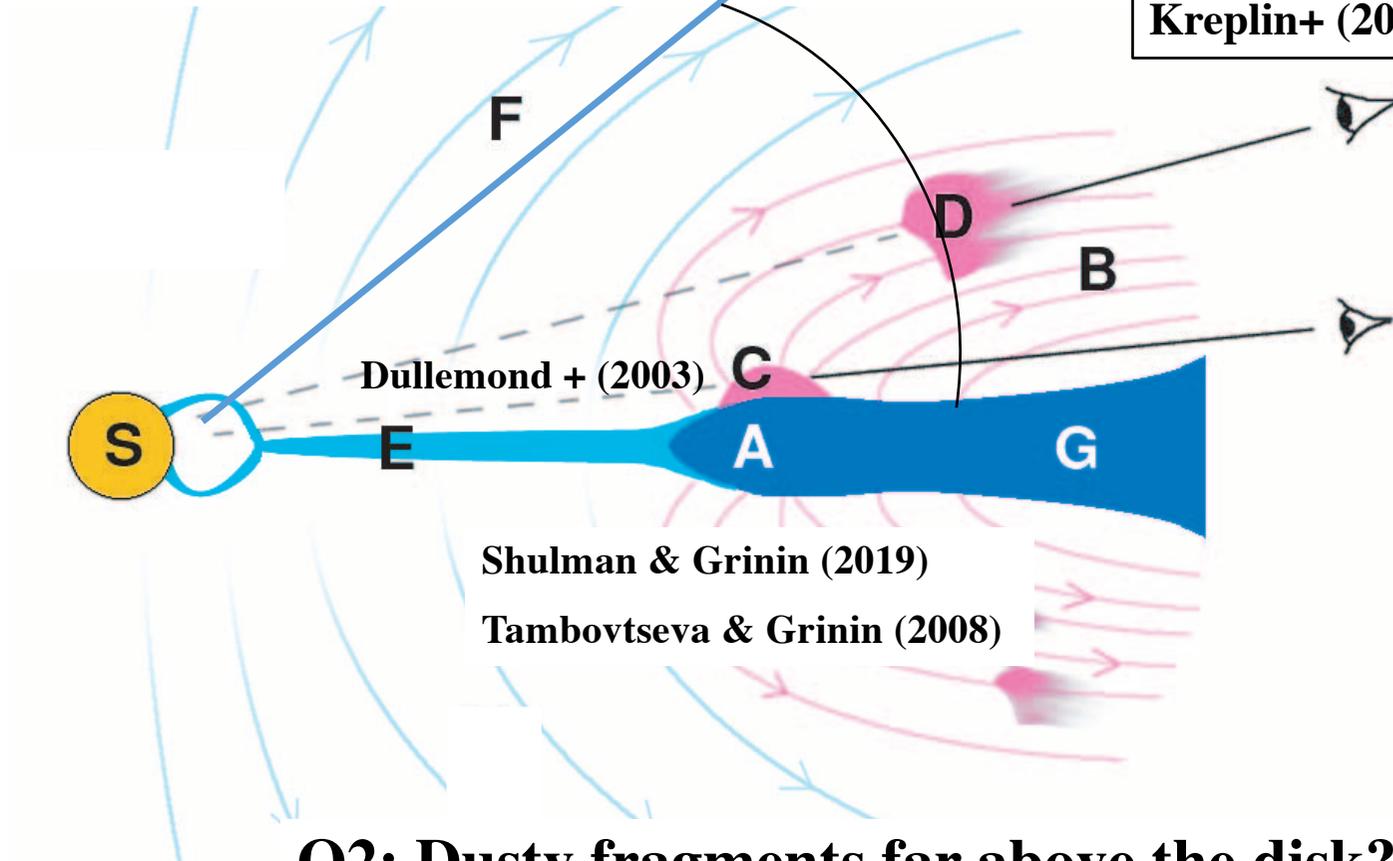
S Cr A (Gahm, Petrov, Tambovtseva, Grinin Stempels, Walter 2018) possible flips < 1 day

Q1: Any colour flips in UXORs?

Vincovic' & Juric' (2007)

> 40° ?

Kreplin+ (2016), Davies + (2018)

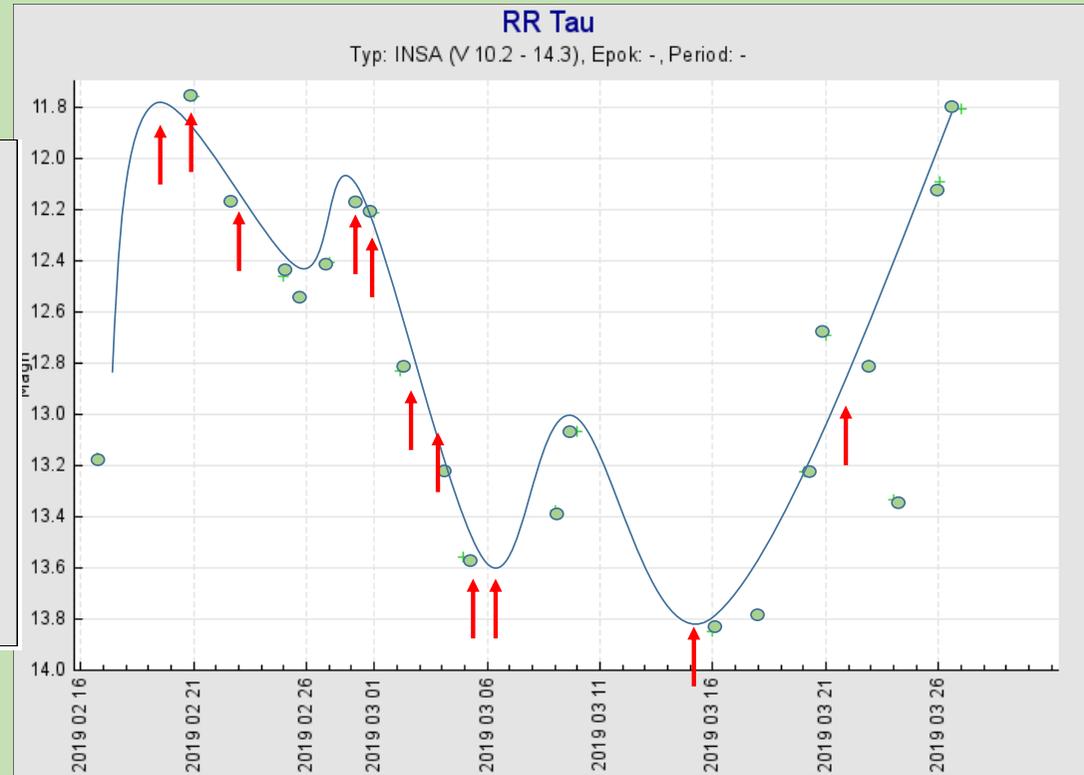
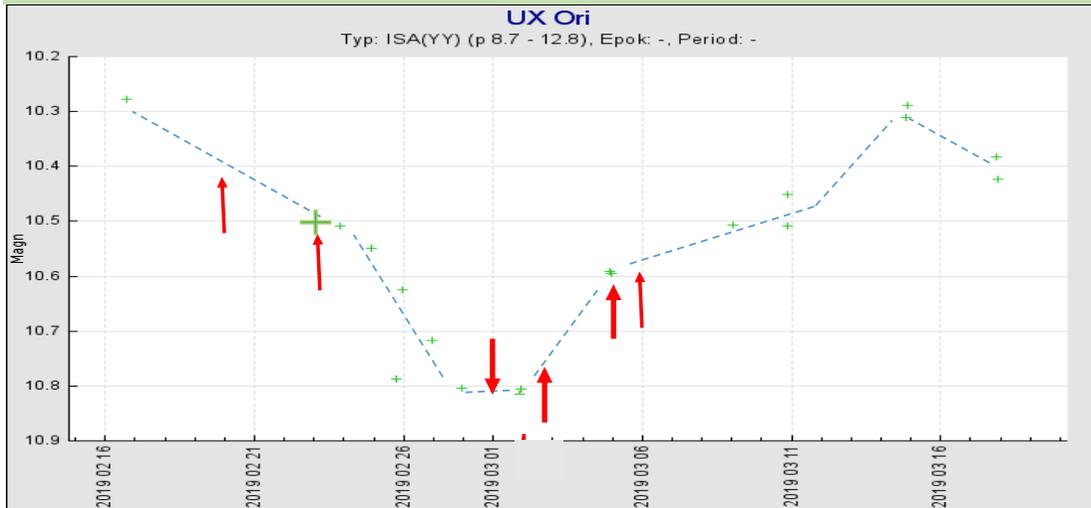


Q2: Dusty fragments far above the disk?

**2019: Gahm, Djupvik (NOT) + Luleå Technical University
also: Grinin and Krelowski**

NOT: spectra R = 25 000, 3600 – 9400Å.

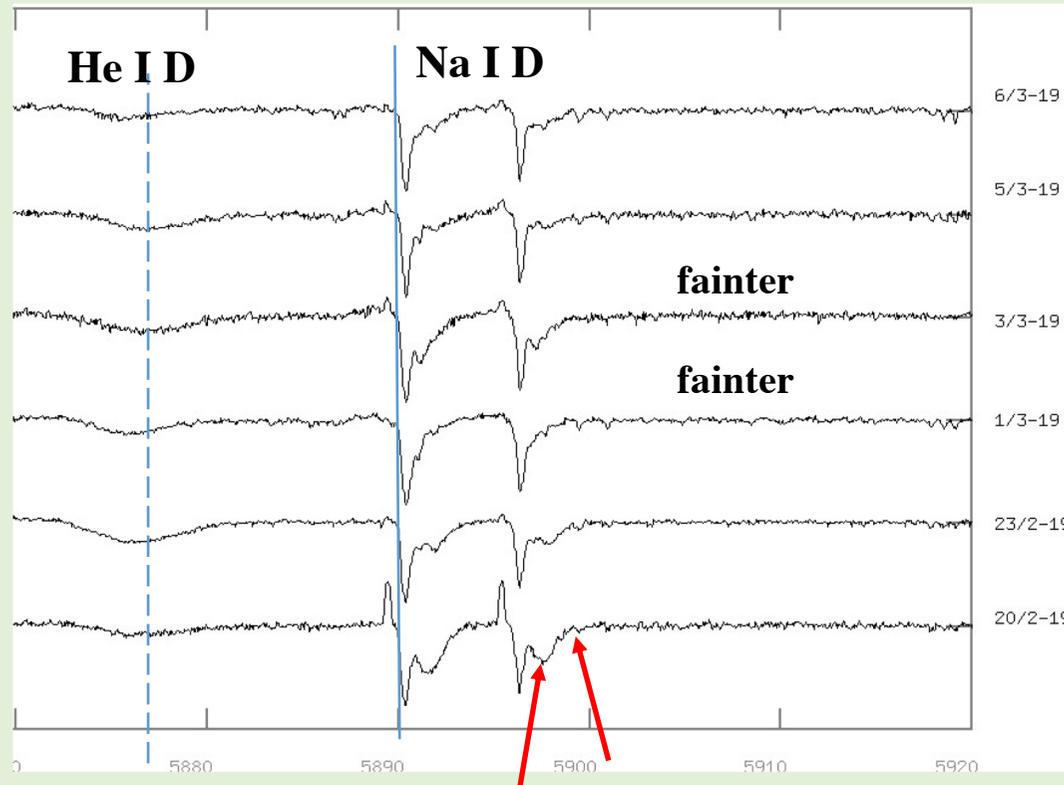
Photometry: Swedish Amateur Astronomy Org.



**Grinin + (2001)
Rodgers + (2002)**

To be continued

UX Ori 2019

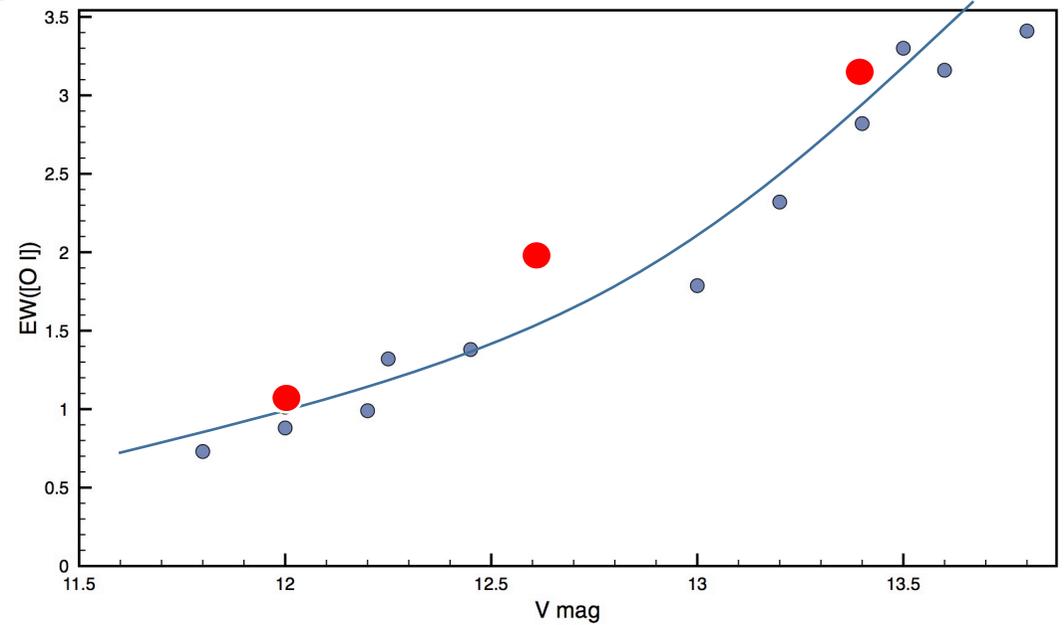
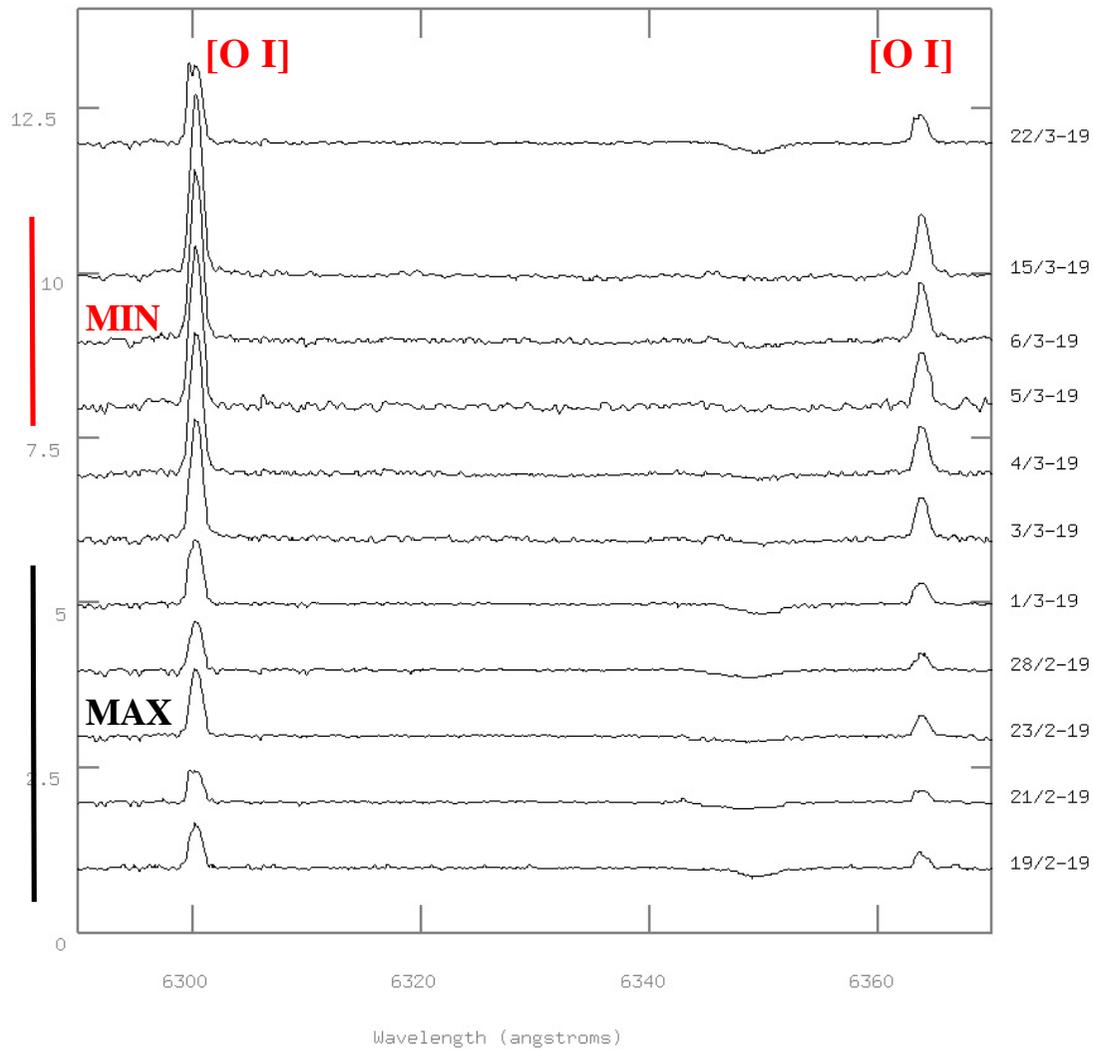


+60 km/s (to +140) Natta+ (2000), Potravnov+ (2019)

During accretion event: no related occultation

Q3: Any evidence of spectroscopic binaries?

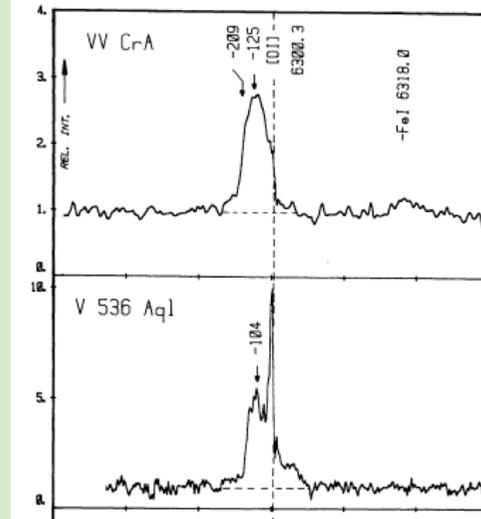
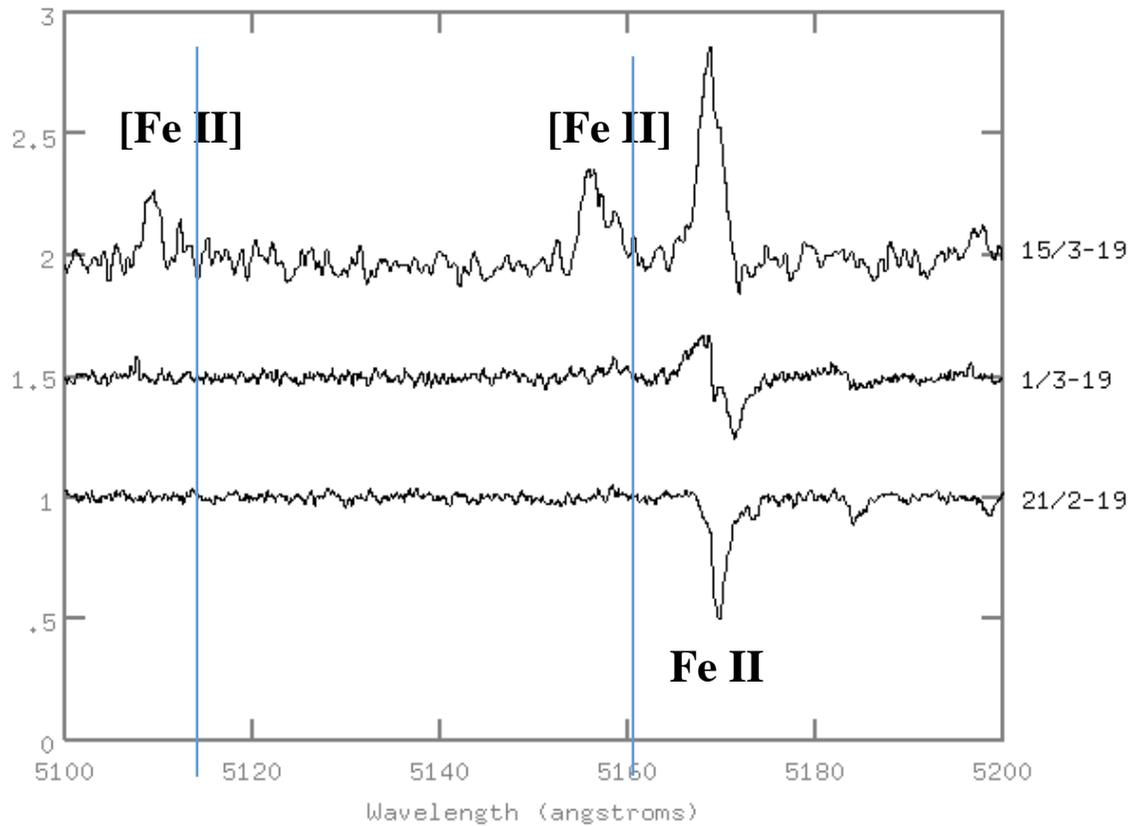
RR Tau 2019



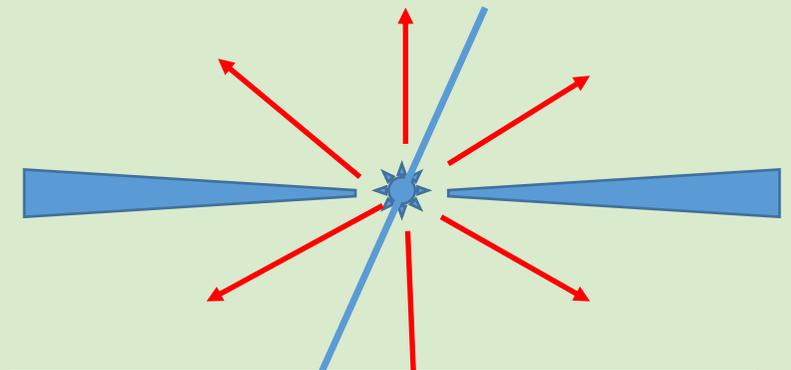
Contrast effect: the line FLUX is very constant

**RV = 0.0 km/s. RV star: +11 km/s ?
 $\Delta V = 58.1$ km/s (FWHM)**

RR Tau 2019



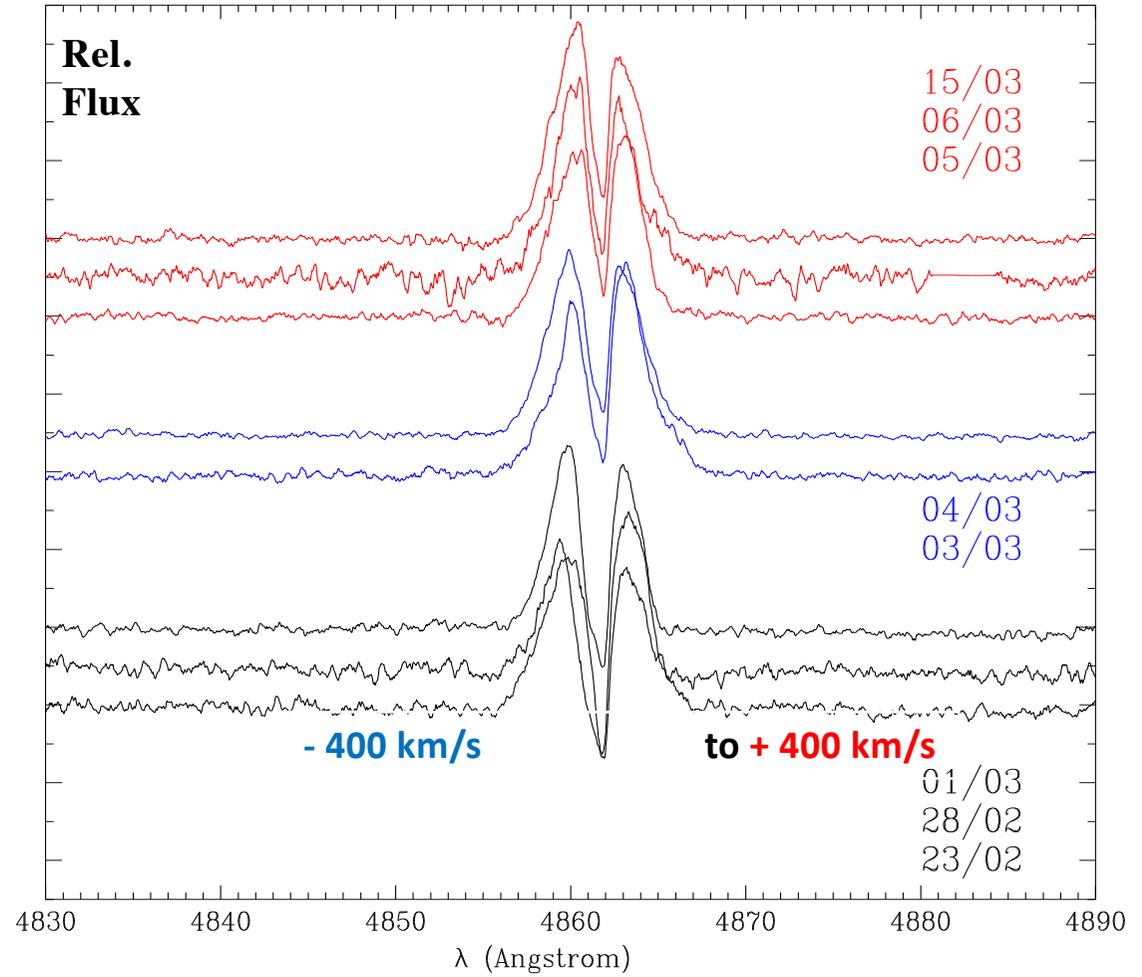
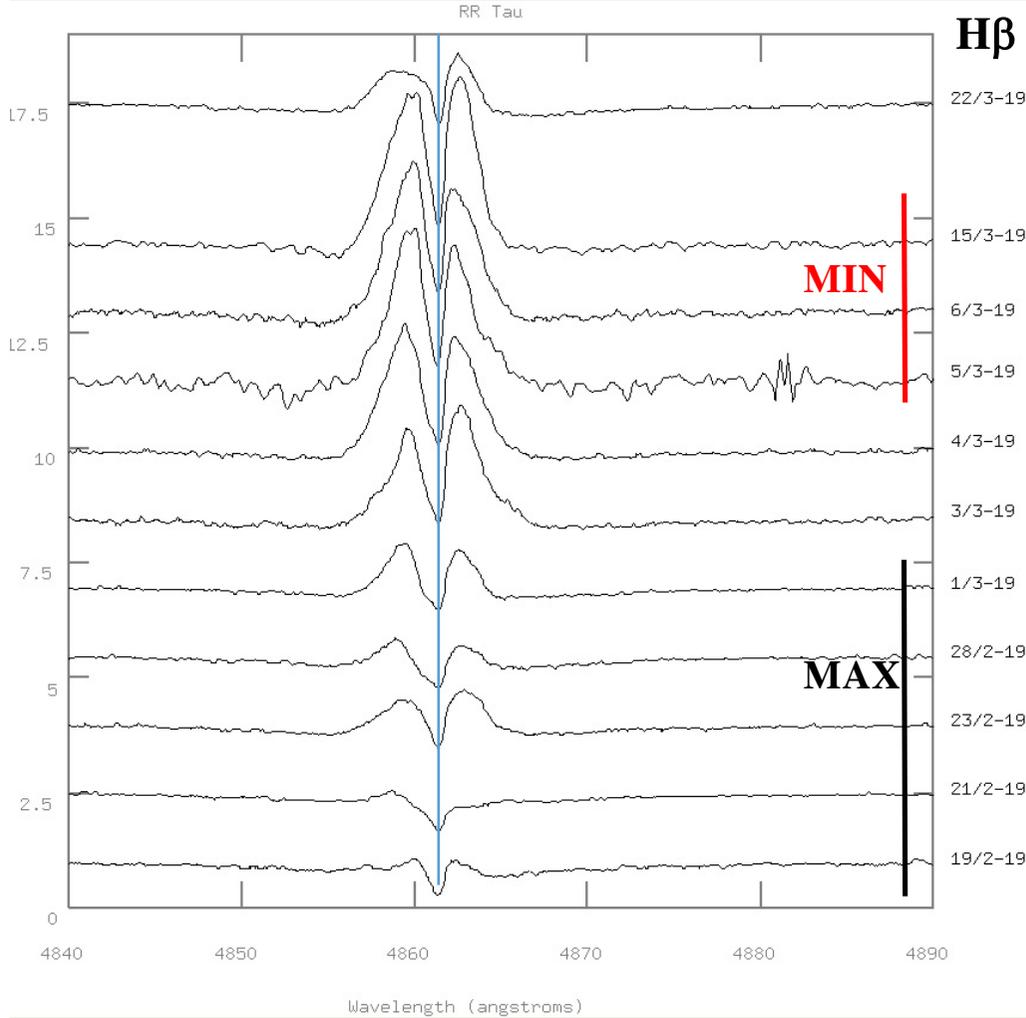
[S II] lines
Appenzeller+ (1984)



[Fe II]: higher critical density than [S II]

Q4: Is the RR Tau disk tilted? Confirmation? ALMA

RR Tau2019



Intrinsic variations: difficult to map inner region

Rodgers et al.

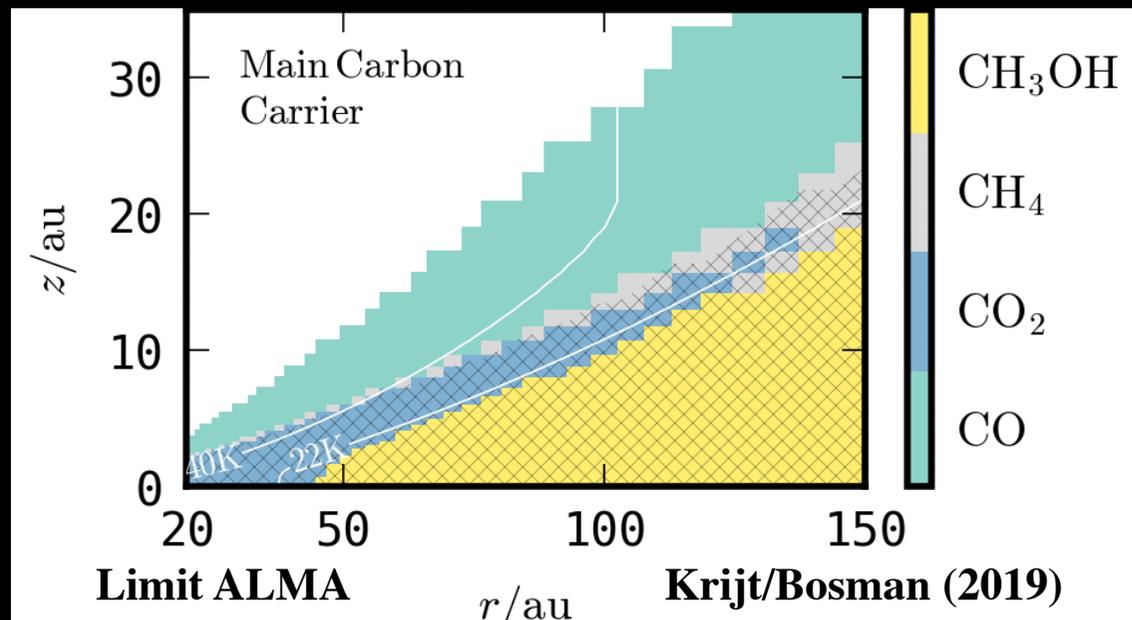
HD 163296

H_2CO

100 AU

MWC 480

100 AU



Q5: Any UXORs with ALMA ?

ANY MOLECULAR LINES IN THE LINE OF SIGHT ?

First test if DIBs are enhanced in UX-objects when occulted:

Andersen, Gahm, Krelowski (1982) for HR 5999

Result: DIBs are not enhanced.

SIGNATURES FROM COLD GAS WHEN RR TAURI IS OCCULTED ?

| <u>Species</u> | <u>Enhanced when occulted?</u> | |
|--|--------------------------------|-------------------------------|
| Ca I, Mg I, K I ground state | no | |
| DIBs | no | (not enhanced in shells/H II) |
| C ₂ (2 – 0), C ₃ | no | (enhanced in shells/H II) |
| CH, CH ⁺ | no | (enhanced in shells/H II) |
| CN | no | |
| NH ₂ | no | |
| Overtone CO 2.3 μ | to be | Ilee+ (2014) |
| Fluorescent H ₂ 2.1 μ | to be | |

Open questions

Q1: Rapid colour flips in UXORs?

Q2: How to bring large bodies with large dust grains far above the disk?

Q3: Evidence of spectroscopic binaries?

Q4: Is the RR Tau disk tilted? Size of inner holes?

Q5: Any mapping of UX Ori stars with ALMA?

Q6: Detecting which molecules in the sightline?